
raytraverse Documentation

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Stephen Wasilewski

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RAYTRAVERSE

raytraverse is a complete workflow for climate based daylight modelling, simulation, and evaluation of architectural spaces. Built around a variance based adaptive sampling strategy, raytraverse can fully explore the daylight conditions throughout a space with efficient use of processing power and storage space.

- Free software: Mozilla Public License 2.0 (MPL 2.0)
- Documentation: <https://raytraverse.readthedocs.io>.

1.1 Installation

The easiest way to install raytraverse is with pip:

```
pip install --upgrade pip setuptools wheel
pip install raytraverse
```

or if you have cloned this repository:

```
cd path/to/this/file
pip install .
```

for a development install (pip install -e may not work correctly):

```
python setup.py develop
```

note that on first run one of the required modules may download some auxiliary data which could take a minute, after that first run start-up is much faster.

1.2 Usage

raytraverse includes a complete command line interface with all commands nested under the *raytraverse* parent command enter:

```
raytraverse --help
```

raytraverse also exposes an object oriented API written primarily in python. calls to Radiance are made through Renderer objects that wrap the radiance c source code in c++ classes, which are made available in python with pybind11. see the src/ directory for more.

For complete documentation of the API and the command line interface either use the Documentation link included above or:

```
pip install -r requirements_dev.txt  
make docs
```

to generate local documentation.

1.3 Git Stuff

this project is hosted in two places, a private repo (master branch) at:

<https://gitlab.enterpriselab.ch/lightfields/raytraverse>

and a public repo (release branch) at:

<https://github.com/stephanwaz/raytraverse>

the repo also depends on two submodules, to initialize run the following:

```
git clone https://github.com/stephanwaz/raytraverse  
cd raytraverse  
git submodule init  
git submodule update --remote  
git -C src/Radiance config core.sparseCheckout true  
cp src/sparse-checkout .git/modules/src/Radiance/info/  
git submodule update --remote --force src/Radiance
```

after a “git pull” make sure you also run:

```
git submodule update
```

to track with the latest commit used by raytraverse.

1.4 Licence

Copyright (c) 2020 Stephen Wasilewski

This Source Code Form is subject to the terms of the Mozilla Public License, v. 2.0. If a copy of the MPL was not distributed with this file, You can obtain one at <http://mozilla.org/MPL/2.0/>.

1.5 Credits

This package was created with Cookiecutter and the [audreyr/cookiecutter-pypackage](#) project template.

1.5.1 raytraverse.scene

Scene

```
class raytraverse.scene.Scene(outdir, scene=None, area=None, reload=True, over-write=False, ptres=1.0, ptro=0.0, pttol=1.0, viewdir=0, 1, 0, viewangle=360, skyres=10.0, maxspec=0.3, **kwargs)
```

Bases: object

container for scene description

Parameters

- **outdir** (*str*) – path to store scene info and output files

- **scene** (*str, optional (required if not reload)*) – space separated list of radiance scene files (no sky) or octree
 - **area** (*str, optional (required if not reload)*) – radiance scene file containing planar geometry of analysis area or a list of points (line per point, space seperated, first 3 columns x, y, z)
 - **reload** (*bool, optional*) – if True attempts to load existing scene files in new instance overrides ‘overwrite’
 - **overwrite** (*bool, optional*) – if True and outdir exists, will overwrite, else raises a FileExistsError
 - **ptres** (*float, optional*) – final spatial resolution in scene geometry units
 - **ptro** (*float, optional*) – angle in degrees counter-clockwise to point grid
 - **pttol** (*float, optional*) – tolerance for point search when using point list for area
 - **viewdir** ((*float, float, float*), *optional*) – vector (x,y,z) view direction (orients UV space)
 - **viewangle** (*float, optional*) – should be 1-180 or 360
 - **skyres** (*float, optional*) – approximate square patch size (sets sun resolution too)
 - **maxspec** (*float, optional*) – maximum specular transmission in scene (used to clip pdf for sun sampling)
- outdir = None**
path to store scene info and output files
Type str
- maxspec = None**
maximum specular transmission in scene
Type float
- reload = None**
try to reload scene files
Type bool
- view = None**
view translation class
Type raytraverse.viewmapper.ViewMapper
- property skyres**
- property scene**
render scene files (octree)
Getter Returns this samplers’s scene file path
Setter Sets this samplers’s scene file path and creates run files
Type str
- pts()**

SunSetter

```
class raytraverse.scene.SunSetter(scene, srct=0.01, skyro=0.0, reload=True, sunres=10.0,
                                    **kwargs)
```

Bases: object

select suns to sample based on sky pdf and scene.

Parameters

- **scene** (`raytraverse.scene.Scene`) – scene class containing geometry, location and analysis plane
- **srct** (`float, optional`) – threshold of sky contribution for determining appropriate srcn
- **skyro** (`float, optional`) – sky rotation (in degrees, ccw)
- **reload** (`bool`) – if True reloads existing sun positions, else always generates new

srct = None

threshold of sky contribution for determining appropriate srcn

Type float

skyro = None

ccw rotation (in degrees) for sky

Type float

scene = None

`raytraverse.scene.Scene`

property sunres

property sun_kd

sun kdtree for directional queries

property suns

holds sun positions

Getter Returns the sun source array

Setter Set the sun source array and write to files

Type np.array

choose_suns()

load_sky_facs()

direct_view()

write_sun(i)

proxy_src(tsuns, tol=10.0)

check if sun directions have matching source in SunSetter

Parameters

- **tsuns** (`np.array`) – (N, 3) array containing sun source vectors to check
- **tol** (`float`) – tolerance (in degrees)

Returns

- `np.array` – (N,) boolean array if sun has a match
- `np.array` – (N,) index to proxy src

_write_suns(sunfile)

write suns to file

Parameters sunfile –

SunSetterLoc

class raytraverse.scene.**SunSetterLoc** (*scene, loc, skyro=0.0, **kwargs*)

Bases: raytraverse.scene.sunsetter.SunSetter

select suns to sample based on sky pdf and scene.

Parameters

- **scene** (raytraverse.scene.Scene) – scene class containing geometry, location and analysis plane
- **loc** (tuple) – lat, lon, tz (in degrees, west is positive)
- **srct** (float, optional) – threshold of sky contribution for determining appropriate srctn
- **skyro** (float, optional) – sky rotation (in degrees, ccw)
- **reload** (bool) – if True reloads existing sun positions, else always generates new

sky = None

raytraverse.scene.SkyInfo

choose_suns ()

SunSetterPositions

class raytraverse.scene.**SunSetterPositions** (*scene, wea, skyro=0.0, **kwargs*)

Bases: raytraverse.scene.sunsetter.SunSetter

select suns to sample based on sky pdf and scene.

Parameters

- **scene** (raytraverse.scene.Scene) – scene class containing geometry, location and analysis plane
- **wea** (str, np.array, optional) – path to sun position file or wea file, or array of sun positions
- **srct** (float, optional) – threshold of sky contribution for determining appropriate srctn
- **skyro** (float, optional) – sky rotation (in degrees, ccw)
- **reload** (bool) – if True reloads existing sun positions, else always generates new

scene = None

raytraverse.scene.Scene

skyro = None

ccw rotation (in degrees) for sky

Type float

property candidates

raytraverse.scene.SkyInfo

choose_suns ()

SkyInfo

```
class raytraverse.scene.SkyInfo(loc, skyro=0.0)
    Bases: object
    sky location data object

    Parameters
        • loc (tuple) – lat, lon, tz (in degrees, west is positive)
        • skyro (float) – sky rotation (in degrees, ccw)

    skyro = None
    ccw rotation (in degrees) for sky

    Type float

    property solarbounds
        read only extent of solar bounds for given location set via loc

        Getter Returns solar bounds
        Type (np.array, np.array)

    property loc
        scene location

        Getter Returns location
        Setter Sets location and self.solarbounds
        Type (float, float, int)

    in_solarbounds(uv, size=0.0)
        for checking if src direction is in solar transit

    Parameters
        • uv (np.array) – source directions
        • size (float) – offset around UV to test

    Returns result – Truth of ray.src within solar transit
    Return type np.array
```

1.5.2 raytraverse.mapper

SpaceMapper

```
class raytraverse.mapper.SpaceMapper(dfile, ptres=1.0, rotation=0.0, tolerance=1.0)
    Bases: object
    translate between world coordinates and normalized UV space

    rotation = None
    ccw rotation (in degrees) for point grid on plane

    Type float

    tolerance = None
    tolerance for point search when using point list for area

    Type float

    ptres = None
    point resolution for area

    Type float
```

property pt_kd
point kdtree for spatial queries built at first use

property sf
bbox scale factor

property ptshape
shape of point grid

property npts
number of points

property bbox
boundary frame for translating between coordinates [[xmin ymin zmin] [xmax ymax zmax]]

Type np.array

_ro_pts (points, rdir=-1)
rotate points

Parameters

- **points** (*np.ndarray*) – world coordinate points of shape (N, 3)
- **rdir** (-1 or 1) –

rotation direction: -1 to rotate from uv space 1 to rotate to uvspace

uv2pt (uv)
convert UV → world

Parameters **uv** (*np.array*) – normalized UV coordinates of shape (N, 2)

Returns **pt** – world xyz coordinates of shape (N, 3)

Return type np.array

pt2uv (xyz)
convert world → UV

Parameters **xyz** (*np.array*) – world xyz coordinates, shape (N, 3)

Returns **uv** – normalized UV coordinates of shape (N, 2)

Return type np.array

idx2pt (idx)

pts ()

in_area (xyz)
check if point is in boundary path

Parameters **xyz** (*np.array*) – uv coordinates, shape (N, 3)

Returns **mask** – boolean array, shape (N,)

Return type np.array

_rad_scene_to_bbox (plane)

SpaceMapperPt

```
class raytraverse.mapper.SpaceMapperPt(dfile, ptres=1.0, rotation=0.0, tolerance=1.0)
    Bases: raytraverse.mapper.spacemapper.SpaceMapper
    translate between world coordinates and normalized UV space

    property sf
        bbox scale factor

    property ptshape
        shape of point grid

    property bbox
        bounding box

    Type np.array of shape (3,2)

    uv2pt (uv)
        convert UV -> world

        Parameters uv (np.array) – normalized UV coordinates of shape (N, 2)
        Returns pt – world xyz coordinates of shape (N, 3)
        Return type np.array

    pt2uv (xyz)
        convert world -> UV

        Parameters xyz (np.array) – world xyz coordinates, shape (N, 3)
        Returns uv – normalized UV coordinates of shape (N, 2)
        Return type np.array

    idx2pt (idx)
    pts()
    in_area (xyz)
        check if point is in boundary path

        Parameters xyz (np.array) – uv coordinates, shape (N, 3)
        Returns mask – boolean array, shape (N,)
        Return type np.array
```

ViewMapper

```
class raytraverse.mapper.ViewMapper(dxyz=0.0, 1.0, 0.0, viewangle=360.0, name='view',
                                    mtxs=None, imtxs=None)
    Bases: object
    translate between world and normalized UV space based on direction and view angle

    Parameters
        • dxyz (tuple, optional) – central view direction
        • viewangle (float, optional) – if < 180, the horizontal and vertical view angle,
          if greater, view becomes 360,180

    property viewangle
        view angle

    property ymtx
        yaw rotation matrix (to standard z-direction y-up)
```

```

property pmtx
    pitch rotation matrix (to standard z-direction y-up)

property bbox
    bounding box of view

Type np.array of shape (2,2)

property sf
    bbox scale factor

property ivm
    viewmapper for opposite view direction (in case of 360 degree view)

property dxyz
    (float, float, float) central view direction

view2world(xyz, i=0)
world2view(xyz, i=0)
xyz2uv(xyz, i=0)
uv2xyz(uv, i=0)
xyz2xy(xyz, i=0)
pixelrays(res, i=0)
ray2pixel(xyz, res, i=0)
pixel2ray(pxy, res, i=0)
pixel2omega(pxy, res)
ctheta(vec, i=0)
radians(vec, i=0)
degrees(vec, i=0)
in_view(vec, i=0, indices=True)

```

SunMapper

```

class raytraverse.mapper.SunMapper(suns)
    Bases: raytraverse.mapper.viewmapper.ViewMapper
    translate between view and normalized UV space

    Parameters suns (np.array) – dx,dy,dz sun positions

```

1.5.3 raytraverse.renderer

Render

```

class raytraverse.renderer.Renderer(rayargs=None,      scene=None,      nproc=None,
                                         **kwargs)
    Bases: object
    virtual renderer class

    initialized = False
    instance = None
    _pyinstance = None
    Engine = None

```

```
name = None
header = ''
arg_prefix = ''
scene = None
classmethod _set_args(args, iot, nproc)
classmethod initialize(args, scene, nproc=None, **kwargs)
classmethod call(rayfile, store=True, outf=None)
classmethod reset()
classmethod reset_instance()
```

RadianceRenderer

```
class raytraverse.renderer.RadianceRenderer(rayargs=None,           scene=None,
                                             nproc=None, **kwargs)
Bases: raytraverse.renderer.Renderer
Virtual class for wrapping c++ Radiance renderer executable classes
returnbytes = False
classmethod update_param(args, nproc=None, iot='ff')
classmethod initialize(args, scene, nproc=None, iot='ff')
classmethod call(rayfile, store=True, outf=None)
classmethod reset()
classmethod reset_instance()
```

Rtrace

```
raytraverse.renderer.Rtrace
alias of raytraverse.renderer.sprenderer.SPRtrace
```

Rcontrib

```
raytraverse.renderer.Rcontrib
alias of raytraverse.renderer.sprenderer.SPRcontrib
```

SPRenderer

```
class raytraverse.renderer.SPRenderers(rayargs=None,      scene=None,      nproc=None,
                                         **kwargs)
Bases: raytraverse.renderer.Renderer
Subprocess renderer class
cleanup = 'rcalc -if3 -of -e $1=.265074126*$1+.670114631*$2+.064811243*$3'
filter_bad_args = [('-Z\\s*', ''), ('-oZ', '-ov')]
classmethod initialize(args, scene, nproc=None, iot='ff')
classmethod call(rayfile, store=True, outf=None, vecs2stdin=True)
```

SPRtrace

```
class raytraverse.renderer.SPRtrace(rayargs=None, scene=None, nproc=None,  
    **kwargs)  
Bases: raytraverse.renderer.sprenderer.SPRenderer  
Engine = 'rtrace'  
name = 'rtrace'  
classmethod load_source(srcname, **kwargs)
```

SPRcontrib

```
class raytraverse.renderer.SPRcontrib(rayargs=None, scene=None, nproc=None,  
    **kwargs)  
Bases: raytraverse.renderer.sprenderer.SPRenderer  
Engine = 'rcontrib'  
name = 'rcontrib'
```

1.5.4 raytraverse.sampler

Sampler

```
class raytraverse.sampler.Sampler(scene, fdres=9, srcn=1, accuracy=1.0, idres=4,  
    stype='generic', srcdef=None, plotp=False, bands=1,  
    engine_args='', nproc=None, **kwargs)
```

Bases: object

base sampling class

To implement a proper-subclass an engine attribute must be set to a renderer instance prior to calling Sampler.__init__. Also, the method sample must be overridden to properly set up arguments for the renderer.call

Parameters

- **scene** (`raytraverse.scene.Scene`) – scene class containing geometry, location and analysis plane
- **fdres** (`int`, *optional*) – final directional resolution given as log2(res)
- **srcn** (`int`, *optional*) – number of sources return per vector by run
- **accuracy** (`float`, *optional*) – parameter to set threshold at sampling level relative to final level threshold (smaller number will increase sampling, default is 1.0)
- **idres** (`int`, *optional*) – initial direction resolution (as log2(res))
- **stype** (`str`, *optional*) – sampler type (prefixes output files)
- **srcdef** (`str`, *optional*) – path or string with source definition to add to scene
- **plotp** (`bool`, *optional*) – show probability distribution plots at each level (first point only)
- **bands** (`int`, *optional*) – number of spectral bands returned by the engine
- **engine_args** (`str`, *optional*) – command line arguments used to initialize engine
- **nproc** (`int`, *optional*) – number of processors to give to the engine, if None, uses os.cpu_count()

engine = None

bands = None
number of spectral bands / channels returned by renderer based on given renderopts (user ensures these agree).

Type int

samplermap = None
mapper to use for sampling

Type func

srcn = None
number of sources return per vector by run

Type int

idres = None
initial direction resolution (as log2(res))

Type int

weights = None
holds weights for self.draw

Type np.array

stype = None
sampler type

Type str

property compiledscene

property idx
index of next level to sample

Type int

property levels
sampling scheme

Getter Returns the sampling scheme

Setter Set the sampling scheme from (ptres, fdres, skres)

Type np.array

property scene
scene information

Getter Returns this sampler's scene

Setter Set this sampler's scene and create octree with source desc

Type *raytraverse.scene.Scene*

sample (vecf)
generic sample function

_uv2xyz (uv, si)
including to allow overriding mapping behavior of daughter classes

sample_idx (pdraws)
generate samples vectors from flat draw indices

Parameters **pdraws** (*np.array*) – flat index positions of samples to generate

Returns

- **si** (*np.array*) – index array of draws matching samps.shape
- **vecs** (*np.array*) – sample vectors

dump_vecs (*vecs*, *si=None*)

save vectors to file

Parameters

- **vecs** (*np.array*) – ray directions to write
- **si** (*np.array*, *optional*) – sample indices

draw()

draw samples based on detail calculated from weights detail is calculated across direction only as it is the most precise dimension

Returns **pdraws** – index array of flattened samples chosen to sample at next level

Return type *np.array*

update_pdf (*si*, *lum*)

update self.weights (which holds values used to calculate pdf)

Parameters

- **si** (*np.array*) – multidimensional indices to update
- **lum** – values to update with

run_callback()

get_scheme()

run()

execute sampler

SCBinSampler

class `raytraverse.sampler.SCBinSampler` (*scene*, *accuracy=1*, *rcopts='ab 7 -ad 60000 -as 30000 -lw 1e-7'*, ***kwargs*)

Bases: `raytraverse.sampler.sampler.Sampler`

sample contributions from the sky hemisphere according to a square grid transformed by shirley-chiu mapping using rcontrib.

Parameters

- **scene** (`raytraverse.scene.Scene`) – scene class containing geometry, location and analysis plane
- **srcn** (*int*, *optional*) – side of square sky resolution

sample (*vecf*)

call rendering engine to sample sky contribution

Parameters **vecf** (*str*) – path of file name with sample vectors shape (N, 6) vectors in binary float format

Returns **lum** – array of shape (N,) to update weights

Return type *np.array*

SunSampler

```
class raytraverse.sampler.SunSampler(scene, suns, plotp=False, **kwargs)
Bases: object
```

factory class for SingleSunSamplers.

Parameters

- **scene** (`raytraverse.scene.Scene`) – scene class containing geometry, location and analysis plane
- **suns** (`raytraverse.sunsetter.SunSetter`) – sun class containing sun locations.

```
viewsampler = None
raytraverse.sampler.SunViewSampler
```

```
sampleargs = None
sampling arguments for SingleSunSampler
```

Type dict

```
reflsampler = None
raytraverse.sampler.SingleSunSampler
```

```
run(view=True, reflection=True)
```

SingleSunSampler

```
class raytraverse.sampler.SingleSunSampler(scene, suns, sidx, speclevel=9, fdres=10,
                                            accuracy=1, rcopts='-ab 6 -ad 3000 -as
                                            1500 -st 0 -ss 16 -aa .1', keepamb=False,
                                            ambcache=True, **kwargs)
```

Bases: raytraverse.sampler.sampler.Sampler

sample contributions from direct suns.

Parameters

- **scene** (`raytraverse.scene.Scene`) – scene class containing geometry, location and analysis plane
- **suns** (`raytraverse.sunsetter.SunSetter`) – sun class containing sun locations.
- **sidx** (`int`) – sun index to sample
- **speclevel** (`int, optional`) – at this sampling level, pdf is made from brightness of sky sampling rather than progressive variance to look for fine scale specular highlights, this should be atleast 1 level from the end and the resolution of this level should be smaller than the size of the source
- **keepamb** (`bool, optional`) – whether to keep ambient files after run, if kept, a successive call will load these ambient files, so care must be taken to not change any parameters
- **ambcache** (`bool, optional`) – whether the rcopts indicate that the calculation will use ambient caching (and thus should write an -af file argument to the engine)

```
slimit = None
```

controls sampling limit in case of limited contribution

Type float

```
specidx = None
```

index of level at which brightness sampling occurs

Type int
sunpos = None
 sun position x,y,z
Type np.array
pdf_from_sky (skyfield, interp=12, rebuild=False, zero=True, filterpts=True)
sample (vecf)
 call rendering engine to sample sky contribution
Parameters **vecf** (str) – path of file name with sample vectors shape (N, 6) vectors in binary float format
Returns **lum** – array of shape (N,) to update weights
Return type np.array
draw()
 draw samples based on detail calculated from weights detail is calculated across direction only as it is the most precise dimension
Returns **pdraws** – index array of flattened samples chosen to sample at next level
Return type np.array

SunViewSampler

class raytraverse.sampler.SunViewSampler (scene, suns, **kwargs)
 Bases: raytraverse.sampler.sampler.Sampler
 sample view rays to direct suns.
 here idres and fdres are sampled on a per sun basis for a view centered on each sun direction with a view angle of .533 degrees (hardcoded in sunmapper class).

Parameters

- **scene** (raytraverse.scene.Scene) – scene class containing geometry, location and analysis plane
- **suns** (raytraverse.sunsetter.SunSetter) – sun class containing sun locations.
- **loadsrd** (bool) – include suns.rad in base scene initialization. if False, self.engine.load_source must be invoked before call.

property levels
 sampling scheme

Getter Returns the sampling scheme

Setter Set the sampling scheme from (ptres, fdres, skres)

Type np.array

check_viz()

sample (vecf)
 call rendering engine to sample direct view rays

Parameters **vecf** (str) – path of file name with sample vectors shape (N, 6) vectors in binary float format

Returns **lum** – array of shape (N,) to update weights

Return type np.array

```
_uv2xyz (uv, si)
    including to allow overriding mapping behavior of daughter classes

draw ()
    draw first level based on sky visibility

run_callback ()
    post sampling, right full resolution (including interpolated values) non zero rays to result file.
```

1.5.5 raytraverse.lightfield

LightField

```
class raytraverse.lightfield.LightField(scene, rebuild=False, prefix='sky', srcn=1, rmraw=False)
Bases: object
container for accessing sampled data
```

Parameters

- **scene** ([raytraverse.scene.Scene](#)) – scene class containing geometry, location and analysis plane
- **rebuild** (*bool, optional*) – build kd-tree even if one exists
- **prefix** (*str, optional*) – prefix of data files to map

```
rebuild = None
force rebuild kd-tree
```

Type bool

```
prefix = None
prefix of data files from sampler (stype)
```

Type str

```
raw_files()
```

```
property vec
direction vector (3,)
```

```
property lum
luminance (srcn,)
```

```
property omega
solid angle (1,)
```

```
outfile (idx)
```

```
items()
```

LightFieldKD

```
class raytraverse.lightfield.LightFieldKD(scene, rebuild=False, prefix='sky', srcn=1,
                                                 rmraw=False)
Bases: raytraverse.lightfield.lightfield.LightField
```

light field with KDtree structures for spatial query

```
static mk_vector_ball (v)
```

```
property d_kd
list of direction kdtrees
```

Getter Returns kd tree structure

Type list of `scipy.spatial.cKDTree`

property scene
scene information

Getter Returns this integrator's scene

Setter Set this integrator's scene

Type `raytraverse.scene.Scene`

raw_files()
get list of files used to build field

_to_mem_map(ar, offset=0)

_get_v1(npts, pref='', ltype=<class 'raytraverse.lightfield.memarraydict.MemArrayDict'>, os0=0, fvrays=False)

_mk_tree(pref='', ltype=<class 'raytraverse.lightfield.memarraydict.MemArrayDict'>, os0=0)

apply_coef(pi, coefs)

add_to_img(img, mask, pi, i, d, coefs=1, vm=None, radius=3)

query_ray(pi, vecs, interp=1)

query_all_pts(vecs, interp=1)

query_ball(pi, vecs, viewangle=180)

_dview(idx, pdirs, mask, res=800, showsample=True)

direct_view(res=800, showsample=True, items=None)
create a summary image of lightfield for each vpt

SCBinField

class `raytraverse.lightfield.SCBinField(scene, rebuild=False, prefix='sky', **kwargs)`
Bases: `raytraverse.lightfield.lightfieldkd.LightFieldKD`
container for accessing sampled data where every ray has a value for each source

apply_coef(pi, coefs)

SunField

class `raytraverse.lightfield.SunField(scene, suns, rebuild=False, rmraw=False)`
Bases: `raytraverse.lightfield.lightfieldkd.LightFieldKD`
container for sun view data

Parameters

- **scene** (`raytraverse.scene.Scene`) – scene class containing geometry, location and analysis plane
- **suns** (`raytraverse.sunsetter.SunSetter`) – sun class containing sun vectors and SunMapper (passed to SunViewField)
- **rebuild** (`bool, optional`) – build kd-tree even if one exists

suns = None
`raytraverse.sunsetter.SunSetter`

view = None
`raytraverse.lightfield.SunViewField`

```
raw_files()
    get list of files used to build field

_mk_tree (pref='', ltype=<class 'list'>, os0=0)

items()

keymap()

add_to_img (img, mask, pi, i, d, coefs=1, vm=None, radius=3)

direct_view (res=200, showsamp=False, items=None)
    create a summary image of lightfield for each vpt
```

SunViewField

class raytraverse.lightfield.SunViewField(*scene*, *suns*, *rebuild*=False, *rmraw*=False)

Bases: raytraverse.lightfield.LightField

container for sun view data

Parameters

- **scene** ([raytraverse.scene.Scene](#)) – scene class containing geometry, location and analysis plane
- **suns** ([raytraverse.sunsetter.SunSetter](#)) – prefix of data files to integrate
- **rebuild** (*bool*, optional) – build kd-tree even if one exists

suns = None

raytraverse.sunsetter.SunSetter

sunmap = None

raytraverse.sunmapper.SunMapper

raw_files()

get list of files used to build field

property raster

individual pixels forming shape of the sun, stored as uv coordinates with basis viewmapper about sun direction and diameter. indexed like vec, lum, and omega

property scene

scene information

Getter Returns this integrator's scene

Setter Set this integrator's scene

Type [raytraverse.scene.Scene](#)

items()

_grp_by_sun (*vecs*, *lums*, *shape*, *pti*, *suni*)

_build_suns (*vecs*, *lums*, *shape*)

loop through points/suns and group adjacent rays

_to_pix (*rxyz*, *atv*, *vm*, *res*)

_smudge (*cnt*, *omegap*, *omegasp*)

hack to ensure equal energy and max luminance)

add_to_img (*img*, *pi*, *sun*, *vm*)

Parameters

- **img** –
- **pi** –

- **sun** –
- **vm** (`raytraverse.mapper.ViewMapper`) –

metric (*psi, v, s, metricfuncs, **kwargs*)

direct_view (*res=2*)
create a summary image of all sun discs from each of vpts

MemArrayDict

class `raytraverse.lightfield.memarraydict.MemArrayDict`

Bases: dict

a dictionary like object that holds arguments for numpy.memmap, the getter returns a view to the array

static _map (*i*)

values () → an object providing a view on D's values

constructors ()

full_array ()

full_constructor ()

index_strides ()

1.5.6 raytraverse.integrator

Integrator

class `raytraverse.integrator.Integrator` (*skyfield, sunfield=None, wea=None, loc=None, skyro=0.0, **kwargs*)

Bases: object

class to generate outputs from skyfield sunfield and sky conditions

This class provides an interface to:

1. generate sky data using the perez
2. combine lightfield data for sky and sun (handling sparsely populated sun data)
3. apply sky data to the lightfield queries
4. output luminance maps and photometric quantities and visual comfort metrics

Parameters

- **skyfield** (`raytraverse.lightfield.SCBinField`) – class containing sky data
- **sunfield** (`raytraverse.lightfield.SunField`) – class containing sun data (should share its scene parameter with the given skyfield)
- **wea** (*str, np.array, optional*) – path to epw, wea, or .npy file or np.array, if loc not set attempts to extract location data (if needed). The Integrator does not need to be initialized with weather data but for convinience can be. If skydata is not set then the optional parameter of get_sky_matrix() is required or an Exception will be raised.
- **loc** (*(float, float, int), optional*) – location data given as lat, lon, mer with + west of prime meridian overrides location data in wea (but not in sunfield)
- **skyro** (*float, optional*) – angle in degrees counter-clockwise to rotate sky (to correct model north, equivalent to clockwise rotation of scene) does not override rotation in SunField)

```
scene = None
    raytraverse.scene.Scene

suns = None
    raytraverse.sunsetter.SunSetter

sky = None
    raytraverse.scene.SkyInfo

sunfield = None
    raytraverse.lightfield.SunField

skyfield = None
    raytraverse.lightfield.SCBinField

property skydata
    sky data formatted as dx, dy, dz, dirnorm, diffhoriz

    Getter Returns this scene's skydata
    Setter Sets this scene's skydata from file path or
    Type np.array

format_skydata(dat)
    process dat argument as skydata

    Parameters dat (str, np.array) – This method takes either a file path or np.array.
        File path can point to a wea, epw, or .npy file. Loaded array must be one of the following:
        - 4 col: alt, az, dir, diff - 5 col: dx, dy, dz, dir, diff - 5 col: m, d, h, dir, diff'

    Returns dx, dy, dz, dir, diff
    Return type np.array

get_sky_mtx(skydata=None)
    generate sky, grnd and sun coefficients from sky data using perez

    Parameters skydata (str, np.array, optional) – if None, uses object skydata
        (will raise error if unassigned) see format_skydata() for argument specifics.

    Returns
        • smtx (np.array) – shape (len(skydata), skyres**2) coefficients for each sky patch each
            row is a timestep, timesteps where a sun exists exclude the sun coefficient, otherwise
            the patch enclosing the sun position contains the energy of the sun
        • grnd (np.array) – shape (len(skydata),) coefficients for ground at each timestep
        • sun (np.array) – shape (len(skydata), 5) sun position (index 0,1,2) and coefficients
            for sun at each timestep assuming the true solid angle of the sun (index 3) and the
            weighted value for the sky patch (index 4).
        • si (np.array) – shape (len(skydata), 2) index for self.suns.suns pointing to correct
            proxy source (col 0) and sunbin for patch mapping (col 1)
        • daysteps (np.array) – shape (len(skydata),) boolean array masking timesteps when
            sun is below horizon

header()
    generate image header string

_hdr(res, vm, keymap, pi, suni, sun, pdirs, si, vstr, outf, interp, mask, skyv)
_hdr(pts, smtx, grnd=None, suns=None, suni=None, daysteps=None, vname='view', viewangle=180.0, res=400, interp=1)

    Parameters
        • pts (np.array) – points and view dirs shape (N, 6)
```

- **smtx** (*np.array*) – shape (len(skydata), skyres**2) coefficients for each sky patch each row is a timestep, timesteps where a sun exists exclude the sun coefficient, otherwise the patch enclosing the sun position contains the energy of the sun
- **grnd** (*np.array*) – shape (len(skydata),) coefficients for ground at each timestep
- **suns** (*np.array*) – shape (len(skydata), 5) sun position (index 0,1,2) and coefficients for sun at each timestep assuming the true solid angle of the sun (index 3) and the weighted value for the sky patch (index 4).
- **suni** (*np.array*) – shape (len(skydata), 2) index for self.suns.suns pointing to correct proxy source (col 0) and sunbin for patch mapping (col 1)
- **daysteps** (*np.array*) – shape (len(skydata),) boolean array masking timesteps when sun is below horizon
- **vname** (*str*) – view name for output file
- **viewangle** (*float, optional*) – degree opening of view cone
- **res** (*int, optional*) – image resolution
- **interp** (*int, optional*) – number of nearest points to interpolate between. 1 will resemble voronoi patches

metric (*pts, smtx, grnd=None, suns=None, suni=None, daysteps=None, metricfuncs=(<function illum>,), **kwargs*)
 calculate luminance based metrics for given sensors and skyvecs

Parameters

- **pts** (*np.array*) – points and view dirs shape (N, 6)
- **smtx** (*np.array*) – shape (len(skydata), skyres^2) coefficients for each sky patch each row is a timestep, timesteps where a sun exists exclude the sun coefficient, otherwise the patch enclosing the sun position contains the energy of the sun
- **grnd** (*np.array*) – shape (len(skydata),) coefficients for ground at each timestep
- **suns** (*np.array*) – shape (len(skydata), 5) sun position (index 0,1,2) and coefficients for sun at each timestep assuming the true solid angle of the sun (index 3) and the weighted value for the sky patch (index 4).
- **suni** (*np.array*) – shape (len(skydata), 2) index for self.suns.suns pointing to correct proxy source (col 0) and sunbin for patch mapping (col 1)
- **daysteps** (*np.array*) – shape (len(skydata),) boolean array masking timesteps when sun is below horizon
- **metricfuncs** (*tuple*) – list of metric functions to apply, with the call signature: f(view, rays, omegas, lum, kwargs)

Returns metrics – metrics at each point/direction and sky weighting axes ordered as: (view, points, skys, metrics) an additional column on the final axis is 0 when only the skyfield is used, 1 when sun reflections exists and 2 when a view to the sun also exists.

Return type np.array

1.5.7 raytraverse.craytraverse

1.5.8 raytraverse.draw

wavelet and associated probability functions.

```
raytraverse.draw.get_detail(samps, axes)
```

run high pass filter over given axes

```
raytraverse.draw.from_pdf(pdf, threshold)
```

1.5.9 raytraverse.io

functions for reading and writing

```
class raytraverse.io.CaptureStdOut(b=False, store=True, outf=None)
```

Bases: object

redirect output streams at system level (including c printf)

Parameters

- **b** (bool, optional) – read data as bytes
- **store** (bool, optional) – record stdout in a IOStream, value accessible through self.stdout
- **outf** (IOBase, optional) – if not None, must be writable, closed on exit

Notes

```
with CaptureStdOut() as capture:  
    do stuff  
capout = capture.stdout
```

when using with pytest include the -s flag or this class has no effect

```
property stdout
```

```
drain_bytes()  
    read stdout as bytes
```

```
drain_str()  
    read stdout as unicode
```

```
raytraverse.io.call_sampler(outf, command, vecs, shape)
```

make subprocess call to sampler given as command, expects rgb value as return for each vec

Parameters

- **outf** (str) – path to write out to
- **command** (str) – command line with executable and options
- **vecs** (np.array) – vectors to pass as stdin to command
- **shape** (tuple) – shape of expected output

Returns lums – of length vectors.shape[0]

Return type np.array

```
raytraverse.io.bytefile2rad(f, shape, slc=Ellipsis, subs='ijk,k>ij', offset=0)
```

```
raytraverse.io.np2bytes(ar, dtype='<f')
```

format ar as bytestring

Parameters

- **ar** (*np.array*) –
- **dtype** (*str*) – argument to pass to *np.dtype()*

Returns**Return type** bytes`raytraverse.io.bytes2np(buf, shape, dtype='<f')`

read ar from bytestring

Parameters

- **buf** (*bytes, str*) –
- **shape** (*tuple*) – array shape
- **dtype** (*str*) – argument to pass to *np.dtype()*

Returns**Return type** np.array`raytraverse.io.bytefile2np(f, shape, dtype='<f')`

read binary data from f

Parameters

- **f** (*IOBase*) – file object to read array from
- **shape** (*tuple*) – array shape
- **dtype** (*str*) – argument to pass to *np.dtype()*

Returns necessary for reconstruction**Return type** ar.shape`raytraverse.io.array2hdr(ar, imgf, header=None)`

write 2d np.array (x,y) to hdr image format

Parameters

- **ar** (*np.array*) –
- **imgf** (*file path to right*) –
- **header** (*list of header lines to append to image header*) –

`raytraverse.io.carray2hdr(ar, imgf, header=None)`

write color channel np.array (3, x, y) to hdr image format

Parameters

- **ar** (*np.array*) –
- **imgf** (*file path to right*) –
- **header** (*list of header lines to append to image header*) –

`raytraverse.io.hdr2array(imgf)`

read np.array from hdr image

Parameters **imgf** (*file path of image*) –**Returns** ar**Return type** np.array`raytraverse.io.rgbe2lum(rgbe)`

convert from Radiance hdr rgbe 4-byte data format to floating point luminance.

Parameters `rgbe` (`np.array`) – r,g,b,e unsigned integers according to: <http://radsite.lbl.gov/radiance/refer/filefmts.pdf>

Returns `lum`

Return type luminance in cd/m²

1.5.10 raytraverse.metric

standardized metric functions

`raytraverse.metric.illum`(*vm, vec, omega, lum, scale=179, **kwargs*)

`raytraverse.metric.avgillum`(*vm, vec, omega, lum, scale=179, area=None, **kwargs*)

`raytraverse.metric.sqllum`(*vm, vec, omega, lum, scale=179, area=None, **kwargs*)

1.5.11 raytraverse.plot

functions for plotting data

`raytraverse.plot.save_img`(*fig, ax, outf, title=None*)

`raytraverse.plot.imshow`(*im, figsize=10, 10, outf=None, **kwargs*)

`raytraverse.plot.mk_img_setup`(*lums, bounds=None, figsize=10, 10, ext=1*)

`raytraverse.plot.set_ang_ticks`(*ax, ext*)

`raytraverse.plot.colormap`(*colors, norm*)

`raytraverse.plot.plot_patches`(*ax, patches, patchargs=None*)

1.5.12 raytraverse.quickplot

functions for plotting data

`raytraverse.quickplot.imshow`(*im, figsize=10, 10, outf=None, **kwargs*)

`raytraverse.quickplot.hist`(*lums, bins='auto', outf=None, **kwargs*)

1.5.13 raytraverse.skycalc

functions for loading sky data and computing sun position

`raytraverse.skycalc.read_epw`(*epw*)

read daylight sky data from epw or wea file

Returns `out` – (month, day, hour, dirnorn, difhoriz)

Return type `np.array`

`raytraverse.skycalc.get_loc_epw`(*epw, name=False*)

get location from epw or wea header

`raytraverse.skycalc.sunpos_utc`(*timesteps, lat, lon, builtin=True*)

Calculate sun position with local time

Calculate sun position (altitude, azimuth) for a particular location (longitude, latitude) for a specific date and time (time is in UTC)

Parameters

- `timesteps` (`np.array(datetime.datetime)`) –
- `lon` (`float`) – longitude in decimals. West is +ve

- **lat** (*float*) – latitude in decimals. North is +ve
- **builtin** (*bool*) – use skyfield builtin timescale

Returns

- (*skyfield.units.Angle*, *skyfield.units.Angle*)
- *altitude and azimuth in degrees*

`raytraverse.skycalc.row_2_datetime64(ts, year=2020)`

`raytraverse.skycalc.datetime64_2_datetime(timesteps, mer=0.0)`
convert datetime representation and offset for timezone

Parameters

- **timesteps** (*np.array(np.datetime64)*) –
- **mer** (*float*) – Meridian of the time zone. West is +ve

Returns

Return type `np.array(datetime.datetime)`

`raytraverse.skycalc.sunpos_degrees(timesteps, lat, lon, mer, builtin=True, ro=0.0)`

Calculate sun position with local time

Calculate sun position (altitude, azimuth) for a particular location (longitude, latitude) for a specific date and time (time is in local time)

Parameters

- **timesteps** (*np.array(np.datetime64)*) –
- **lon** (*float*) – longitude in decimals. West is +ve
- **lat** (*float*) – latitude in decimals. North is +ve
- **mer** (*float*) – Meridian of the time zone. West is +ve
- **builtin** (*bool, optional*) – use skyfield builtin timescale
- **ro** (*float, optional*) – ccw rotation (project to true north) in degrees

Returns Sun position as (altitude, azimuth) in degrees

Return type `np.array([float, float])`

`raytraverse.skycalc.sunpos_radians(timesteps, lat, lon, mer, builtin=True, ro=0.0)`

Calculate sun position with local time

Calculate sun position (altitude, azimuth) for a particular location (longitude, latitude) for a specific date and time (time is in local time)

Parameters

- **timesteps** (*np.array(np.datetime64)*) –
- **lon** (*float*) – longitude in decimals. West is +ve
- **lat** (*float*) – latitude in decimals. North is +ve
- **mer** (*float*) – Meridian of the time zone. West is +ve
- **builtin** (*bool*) – use skyfield builtin timescale
- **ro** (*float, optional*) – ccw rotation (project to true north) in radians

Returns Sun position as (altitude, azimuth) in radians

Return type `np.array([float, float])`

`raytraverse.skycalc.sunpos_xyz (timesteps, lat, lon, mer, builtin=True, ro=0.0)`

Calculate sun position with local time

Calculate sun position (altitude, azimuth) for a particular location (longitude, latitude) for a specific date and time (time is in local time)

Parameters

- **timesteps** (`np.array (np.datetime64)`) –
- **lon** (`float`) – longitude in decimals. West is +ve
- **lat** (`float`) – latitude in decimals. North is +ve
- **mer** (`float`) – Meridian of the time zone. West is +ve
- **builtin** (`bool`) – use skyfield builtin timescale
- **ro** (`float, optional`) – ccw rotation (project to true north) in degrees

Returns Sun position as (x, y, z)

Return type `np.array`

`raytraverse.skycalc.generate_wea (ts, wea, interp='linear')`

`raytraverse.skycalc.coeff_lum_perez (sunz, epsilon, delta, catn)`

matches coeff_lum_perez in gendaylit.c

`raytraverse.skycalc.perez_apply_coef (coefs, cgamma, dz)`

`raytraverse.skycalc.perez_lum_raw (tp, dz, sunz, coefs)`

matches calc_rel_lum_perez in gendaylit.c

`raytraverse.skycalc.perez_lum (xyz, coefs)`

matches perezlum.cal

`raytraverse.skycalc.perez (sxyz, dirdif, md=None)`

compute perez coefficients

Notes

to match the results of gendaylit, for a given sun angle without associated date, the assumed eccentricity is 1.035020

Parameters

- **sxyz** (`np.array`) – (N, 3) dx, dy, dz sun position
- **dirdif** (`np.array`) – (N, 2) direct normal, diffuse horizontal W/m²
- **md** (`np.array, optional`) – (N, 2) month day of sky calcs (for more precise eccentricity calc)

Returns `perez` – (N, 10) diffuse normalization, ground brightness, perez coeffs, x, y, z

Return type `np.array`

`raytraverse.skycalc.sky_mtx (sxyz, dirdif, side, jn=4)`

generate sky, ground and sun values from sun position and sky values

Parameters

- **sxyz** (`np.array`) – sun directions (N, 3)
- **dirdif** (`np.array`) – direct normal and diffuse horizontal radiation (W/m²) (N, 2)
- **side** (`int`) – sky subdivision
- **jn** (`int`) – sky patch subdivision n = jn²

Returns

- **skymtx** (*np.array*) – (N, side*side)
- **grndval** (*np.array*) – (N,)
- **sunval** (*np.array*) – (N, 4) - sun direction and radiance

1.5.14 raytraverse.translate

functions for translating between coordinate spaces and resolutions

```
raytraverse.translate.norm(v)
    normalize 2D array of vectors along last dimension

raytraverse.translate.norm1(v)
    normalize flat vector

raytraverse.translate.tpnorm(thetaphi)
    normalize angular vector to 0-pi, 0-2pi

raytraverse.translate.uv2xy(uv)
    translate from unit square (0,1),(0,1) to disk (x,y) http://psgraphics.blogspot.com/2011/01/improved-code-for-concentric-map.html

raytraverse.translate.uv2xyz(uv, axes=0, 1, 2, xsign=-1)
    translate from 2 x unit square (0,2),(0,1) to unit sphere (x,y,z) http://psgraphics.blogspot.com/2011/01/improved-code-for-concentric-map.html

raytraverse.translate.xyz2uv(xyz, normalize=False, axes=0, 1, 2, flipu=True)
    translate from vector x,y,z (normalized) to u,v (0,2),(0,1) Shirley, Peter, and Kenneth Chiu. A Low Distortion Map Between Disk and Square. Journal of Graphics Tools, vol. 2, no. 3, Jan. 1997, pp. 45-52. Taylor and Francis+NEJM, doi:10.1080/10867651.1997.10487479.

raytraverse.translate.xyz2xy(xyz, axes=0, 1, 2, flip=True)
raytraverse.translate.pxy2xyz(px, viewangle=180.0)
raytraverse.translate.tp2xyz(thetaphi, normalize=True)
    calculate x,y,z vector from theta (0-pi) and phi (0-2pi) RHS Z-up

raytraverse.translate.xyz2tp(xyz)
    calculate theta (0-pi), phi from x,y,z RHS Z-up

raytraverse.translate.tp2uv(thetaphi)
    calculate UV from theta (0-pi), phi

raytraverse.translate.uv2tp(uv)
    calculate theta (0-pi), phi from UV

raytraverse.translate.uv2ij(uv, side)
raytraverse.translate.uv2bin(uv, side)
raytraverse.translate.bin2uv(bn, side)
raytraverse.translate.bin_borders(sb, side)

raytraverse.translate.resample(samps, ts=None, gauss=True, radius=None)
    simple array resampling. requires whole number multiple scaling.
```

Parameters

- **samps** (*np.array*) – array to resample along each axis
- **ts** (*tuple, optional*) – shape of output array, should be multiple of samps.shape
- **gauss** (*bool, optional*) – apply gaussian filter to upsampling
- **radius** (*float, optional*) – when gauss is True, filter radius, default is the scale ratio - 1

Returns to resampled array

Return type np.array

```
raytraverse.translate.interpolate2d(a, s)
```

```
raytraverse.translate.rmtx_elem(theta, axis=2, degrees=True)
```

```
raytraverse.translate.rotate_elem(v, theta, axis=2, degrees=True)
```

```
raytraverse.translate.rmtx_yp(v)
```

generate a pair of rotation matrices to transform from vector v to z, enforcing a z-up in the source space and a y-up in the destination. If v is z, returns pair of identity matrices, if v is -z returns pair of 180 degree rotation matrices.

Parameters `v`(array-like of size (3,)) – the vector direction representing the starting coordinate space

Returns `ymtx`, `pmtx` – two rotation matrices to be premultiplied in order to reverse transorm, swap order and transpose. Forward: `pmtx@(ymtx@xyz.T)).T` Backward: `ymtx.T@(pmtx.T@xyz.T)).T`

Return type (np.array, np.array)

```
raytraverse.translate.chord2theta(c)
```

compute angle from chord on unit circle

Parameters `c`(float) – chord or euclidean distance between normalized direction vectors

Returns `theta` – angle captured by chord

Return type float

```
raytraverse.translate.theta2chord(theta)
```

compute chord length on unit sphere from angle

Parameters `theta`(float) – angle

Returns `c` – chord or euclidean distance between normalized direction vectors

Return type float

```
raytraverse.translate.aa2xyz(aa)
```

```
raytraverse.translate.xyz2aa(xyz)
```

1.5.15 raytraverse

```
raytraverse [OPTIONS] OUT COMMAND1 [ARGS]... [COMMAND2 [ARGS]...]...
```

the raytraverse executable is a command line interface to the raytraverse python package for running and evaluating climate based daylight models. sub commands of raytraverse can be chained but should be invoked in the order given.

the easiest way to manage options and sure that Scene and SunSetter classes are properly reloaded is to use a configuration file, to make a template:

```
raytraverse --template > run.cfg
```

after adjusting the settings, than each command can be invoked in turn and any dependencies will be loaded with the correct options, a complete run and evaluation can then be called by:

```
raytraverse -c run.cfg OUT sky sunrun integrate
```

as both scene and sun will be invoked automatically as needed.

Arguments:

- `out`: path to new or existing directory for raytraverse run

Arguments

`OUT`

Required argument

Options

VALUE OPTIONS:

`-c, --config <PATH>`
path of config file to load

FLAGS (DEFAULT FALSE):

`--template, --no-template`
write default options to std out as config
Default False

HELP:

`-opts, --opts`
check parsed options
Default False
`--debug`
show traceback on exceptions
Default False
`--version`
Show the version and exit.
Default False

Commands

scene
The scene command creates a Scene object...

sky
the sky command initializes and runs a sky...

suns
the suns command provides a number of options...

sunrun
the sunrun command initializes and runs a...

integrate
the integrate command combines sky and sun...

scene

```
raytraverse scene [OPTIONS]
```

The scene command creates a Scene object which holds geometric information about the model including object geometry (and defined materials), the analysis plane and the desired resolutions for sky and analysis plane subdivision

Options

VALUE OPTIONS:

-area <TEXT>

radiance scene file containing planar geometry of analysis area

-maxspec <FLOAT>

an important parameter for guiding reflected sun rays. contribution values above this threshold are assumed to be direct view rays. If possible, (1) this value should be less than the tvis of the darkest glass in the scene, and (2) greater than the highest expected contribution from a specular reflection or scattering interaction. If it is not possible to meet both conditions, then ensure that condition (2) is met and consider using a substantially higher skyres to avoid massive over sampling of direct view rays

Default 0.3

-ptres <FLOAT>

resolution of point subdivision on analysis plane. units match radiance scene file

Default 2.0

-scene <TEXT>

space separated list of radiance scene files (no sky) or precompiled octree

-skyres <FLOAT>

sky is subdivided according to a shirley-chiu disk to square mapping, total number of sky bins will equal skyres². solid angle of each patch will be 2*pi/(skyres²)

Default 10.0

FLAGS (DEFAULT TRUE):

--reload, --no-reload

if a scene already exists at OUT reload it, note that if this is False and overwrite is False, the program will abort

Default True

FLAGS (DEFAULT FALSE):

--info, --no-info

print info on scene to stderr

Default False

--overwrite, --no-overwrite

Warning! if set to True and reload is False all files in OUT will be deleted

Default False

--points, --no-points

print point locations to stdout

Default False

HELP:**-opts, --opts**

check parsed options

Default False**--debug**

show traceback on exceptions

Default False**--version**

Show the version and exit.

Default False**sky**

```
raytraverse sky [OPTIONS]
```

the sky command initializes and runs a sky sampler and then readies the results for integration by building a SCBinField. sky should be invoked before calling suns, as the sky contributions are used to select the necessary sun positions to run

Options**VALUE OPTIONS:****-accuracy <FLOAT>**

a generic accuracy parameter that sets the threshold variance to sample. A value of 1 will have a sample count at the final sampling level equal to the number of directions with a contribution variance greater than .25

Default 1.0**-fdres <INTEGER>**

the final directional sampling resolution, yielding a grid of potential samples at $2^{fdres} \times 2^{fdres}$ per hemisphere

Default 9**-idres <INTEGER>**

the initial directional sampling resolution. each side of the sampling square (representing a hemisphere) will be subdivided 2^{idres} , yielding $2^{(2*idres)}$ samples and a resolution of $2^{(2*idres)/(2\pi)}$ samples/steradian. this value should be smaller than 1/2 the size of the smallest view to an aperture that should be captured with 100% certainty

Default 4**-rcopts <TEXT>**

rtrace options to pass to the rcontrib call see the man pages for rtrace, rcontrib, and rcontrib -defaults for more information

Default -ab 7 -ad 60000 -as 30000 -lw 1e-7 -st 0 -ss 16

FLAGS (DEFAULT TRUE):

--rmraw, --no-rmraw

if True removes output of sampler.run(), after SCBinField is constructed. Note that SCBinField cannot be rebuilt once raw files are removed

Default True

--run, --no-run

if True calls sampler.run()

Default True

FLAGS (DEFAULT FALSE):

--plotdview, --no-plotdview

plot a direct view of the sky field (as a .hdr file), this is equivalent to integrating with a value of 1 for all sky patches with no interpolation, plots pixels of actualsample vectors in red

Default False

--plotp, --no-plotp

for diagnostics only, plots the pdf at each level for point[0,0] in an interactive display (note that program will hang until the user closes the plot window at each level)

Default False

HELP:

-opts, --opts

check parsed options

Default False

--debug

show traceback on exceptions

Default False

--version

Show the version and exit.

Default False

suns

```
raytraverse suns [OPTIONS]
```

the suns command provides a number of options for creating sun positions used by sunrun see wea and usepositions options for details

Note:

the wea and skyro parameters are used to reduce the number of suns in cases where a specific site is known. Only suns within the solar transit (or positions if usepositions is True will be selected. It is important to note that when integrating, if a sun position outside this range is queried than results will not include the more detailed simulations involved in sunrun and will instead place the suns energy within the nearest sky patch. if skyres is small and or the patch is directly visible this will introduce significant bias in most metrics.

Options

VALUE OPTIONS:

-loc <FLOATS>

specify the scene location (if not specified in -wea or to override, give as “lat lon mer” where lat is + North, lon is + West and mer is the timezone meridian (full hours are 15 degree increments)

-skyro <FLOAT>

counter clockwise rotation (in degrees) of the sky to rotate true North to project North, so if project North is 10 degrees East of North, skyro=10

Default 0.0

-srct <FLOAT>

if the contribution of a sky patch (for any view ray) is above this threshold, a sun will be created in this patch

Default 0.01

-sunres <FLOAT>

resolution in degrees of the sky patch grid in which to stratify sun samples. Suns are randomly located within the grid, so this corresponds to the average distance between sources. The average error to a randomly selected sun position will be on average ~0.4 times this value

Default 10.0

-wea <TEXT>

path to weather/sun position file. possible formats are:

1. .wea file
2. .wea file without header (require -loc and –no-usepositions)
3. .epw file
4. .epw file without header (require -loc and –no-usepositions)
5. 3 column tsv file, each row is dx, dy, dz of candidate sun position (requires –usepositions)
6. 4 column tsv file, each row is altitude, azimuth, direct normal, diff. horizontal of candidate suns (requires –usepositions)
7. 5 column tsv file, each row is dx, dy, dz, direct normal, diff. horizontal of candidate suns (requires –usepositions)

tsv files are loaded with loadtxt

FLAGS (DEFAULT TRUE):

--reload, --no-reload

if False, regenerates sun positions, because positions may be randomly selected this will make any sunrun results obsolete

Default True

FLAGS (DEFAULT FALSE):

--plotdview, --no-plotdview

creates a png showing sun positions on an angular fisheye projection of the sky. sky patches are colored by the maximum contributing ray to the scene

Default False

--usepositions, --no-usepositions

if True, sun positions will be chosen from the positions listed in wea. if more than one position is a candidate for that particular sky patch (as determined by sunres) than a random choice will be made. by using one of the tsv format options for wea, and preselecting sun positions such that there is 1 per patch a deterministic result canbe achieved.

Default False

HELP:

-opts, --opts

check parsed options

Default False

--debug

show traceback on exceptions

Default False

--version

Show the version and exit.

Default False

sunrun

```
raytraverse sunrun [OPTIONS]
```

the sunrun command initializes and runs a sun sampler and then readies the results for integration by building a SunField.

Options

VALUE OPTIONS:

-accuracy <FLOAT>

a generic accuracy parameter that sets the threshold variance to sample. A value of 1 will have a sample count at the final sampling level equal to the number of directions with a contribution variance greater than .25

Default 1.0

-fdres <INTEGER>

the final directional sampling resolution, yielding a grid of potential samples at $2^{fdres} \times 2^{fdres}$ per hemisphere

Default 10

-idres <INTEGER>

the initial directional sampling resolution. each side of the sampling square (representing a hemisphere) will be subdivided 2^{idres} , yielding $2^{(2*idres)}$ samples and a resolution of $2^{(2*idres)/(2\pi)}$ samples/steradian. this value should be smaller than 1/2 the size of the smallest view to an aperture that should be captured with 100% certainty

Default 4

-rcopts <TEXT>

rtrace options for sun reflection runs see the man pages for rtrace, and rtrace -defaults for more information

Default -ab 6 -ad 3000 -as 1500 -st 0 -ss 16 -aa .1

-speclevel <INTEGER>

at this sampling level, pdf is made from brightness of sky sampling rather than progressive variance to look for fine scale specular highlights, this should be atleast 1 level from the end and the resolution of this level should be smaller than the size of the source

Default 9

FLAGS (DEFAULT TRUE):

--ambcache, --no-ambcache

whether the rcopts indicate that the calculation will use ambient caching (and thus should write an -af file argument to the engine)

Default True

--reflection, --no-reflection

run/build/plot reflected sun components

Default True

--rmraw, --no-rmraw

if True removes output of sampler.run(), after SCBinField is constructed. Note that SCBinField cannot be rebuilt once raw files are removed

Default True

--run, --no-run

if True calls sampler.run()

Default True

--view, --no-view

run/build/plot direct sun views

Default True

FLAGS (DEFAULT FALSE):

--keepamb, --no-keepamb

whether to keep ambient files after run, if kept, a successive call will load these ambient files, so care must be taken to not change any parameters

Default False

--plotdview, --no-plotdview

plot a direct view of the sky field (as a .hdr file), this is equivalent to integrating with a value of 1 for all sky patches with no interpolation, plots pixels of actualsample vectors in red

Default False

--plotp, --no-plotp

for diagnostics only, plots the pdf at each level for point[0,0] in an interactive display (note that program will hang until the user closes the plot window at each level)

Default False

HELP:

-opts, --opts

check parsed options

Default False

--debug

show traceback on exceptions

Default False

--version

Show the version and exit.

Default False

integrate

```
raytraverse integrate [OPTIONS]
```

the integrate command combines sky and sun results and evaluates the given set of positions and sky conditions

Options

VALUE OPTIONS:

-interp <INTEGER>

number of nearby rays to use for interpolation of hdrroutput (weighted by a gaussian filter). this doesnot apply to metric calculations

Default 12

-loc <FLOATS>

specify the scene location (if not specified in -wea or to override. give as “lat lon mer” where lat is + North, lon is + West and mer is the timezone meridian (full hours are 15 degree increments)

-pts <TEXT>

points to evaluate, this can be a .npy file, a whitespace seperated text file or entered as a string with commas between components of a point and spaces between points. in all cases each point requires 6 numbers x,y,z,dx,dy,dz so the shape of the array will be (N, 6)

Default 0,0,0,-1,0

-res <INTEGER>

the resolution of hdr output in pixels

Default 800

-skyro <FLOAT>

counter clockwise rotation (in degrees) of the sky to rotate true North to project North, so if project North is 10 degrees East of North, skyro=10

Default 0.0

-vname <TEXT>

name to include with hdr outputs

Default view

-wea <TEXT>

path to weather/sun position file. possible formats are:

1. .wea file

2. .wea file without header (requires -loc)
 3. .epw file
 4. .epw file without header (requires -loc)
 5. 4 column tsv file, each row is altitude, azimuth, direct normal, diff. horizontal of candidate suns (requires --usepositions)
 6. 5 column tsv file, each row is dx, dy, dz, direct normal, diff. horizontal of candidate suns (requires --usepositions)
- tsv files are loaded with loadtxt

FLAGS (DEFAULT TRUE):

--hdr, --no-hdr
produce an hdr output for each point and line in wea

Default True

--metric, --no-metric
calculate metrics for each point and wea file output is ordered by point than sky

Default True

FLAGS (DEFAULT FALSE):

--header, --no-header
print column headings on metric output

Default False

--skyonly, --no-skyonly
if True, only integrate on Sky Field, useful for diagnostics

Default False

HELP:

-opts, --opts
check parsed options

Default False

--debug
show traceback on exceptions

Default False

--version
Show the version and exit.

Default False

1.5.16 History

0.2.0 (2020-09-25)

- Build now includes all radiance dependencies to setup multi-platform testing
- In the absence of craytraverse, sampler falls back to SPRenderer
- install process streamlined for developer mode
- travis ci deploys linux and mac wheels directly to pypi
- **release.sh should be run after updating this file, tests past locally and docs build.**

0.1.0 (2020-05-19)

- First release on PyPI.

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1.5.18 Search

1.5.19 Todo

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